



The formation rate and luminosity function of fast radio bursts

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Summary

- Sample: non-repeating FRBs from the first CHIME/FRB catalog.
- Method: Lynden-Bell's c^- method.
- Result
 - A relatively strong luminosity evolution. The luminosity function of FRBs can be well fitted with a broken power-law model.
 - The formation rate decreases rapidly, with similar to the GRBs.
 - FRBs are associated with older stellar populations.

Introduction

● what is the source of FRBs?

- old neutron stars
- binary neutron star (white dwarf) mergers
- neutron star Colliding/black hole?
- Exploding/colliding stars?
- Colliding neutron stars?
- Bursting magnetars?
- Comets/asteroids impacting neutron stars?
- Evaporating black holes?

Introduction

- **Observationally**

- **FRBs: repeaters and non-repeaters.**

- Some FRBs are produced by magnetars (Death of massive stars), eg., FRB 20200428, **FRBs follow the star formation history (SFR).**

- Some FRBs are related to ancient star populations, eg., FRB 20200120E, **FRBs do not follow the SFR.**

Research questions: **Unveiling the Physical Origin of FRBs.**

Introduction

●Related Methods

Methods



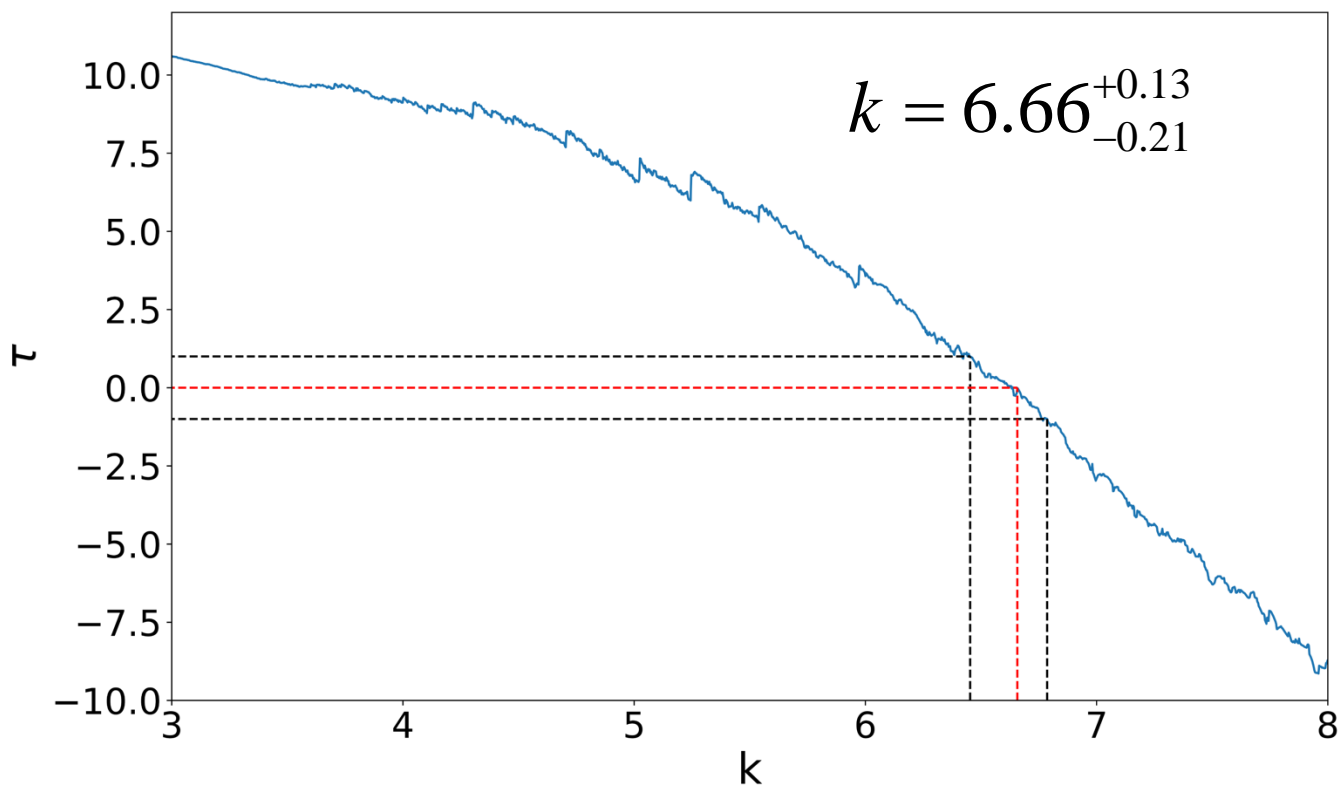
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graph LR; Methods[Methods] --- Box1[FRB formation rate vs SFR<br/>(Zhang & Wang (2019))]; Methods --- Box2[The number density of FRB sources vs the density of<br/>possible ancestors]; Methods --- Box3[Redshift evolution of the luminosity or energy function<br/>of FRBs.<br/>(James et al. 2022; Hashimoto et al. 2022)]
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FRB formation rate vs SFR_{(Zhang & Wang (2019))}

The number density of FRB sources vs the density of possible ancestors

Redshift evolution of the luminosity or energy function of FRBs._(James et al. 2022; Hashimoto et al. 2022)

Result:

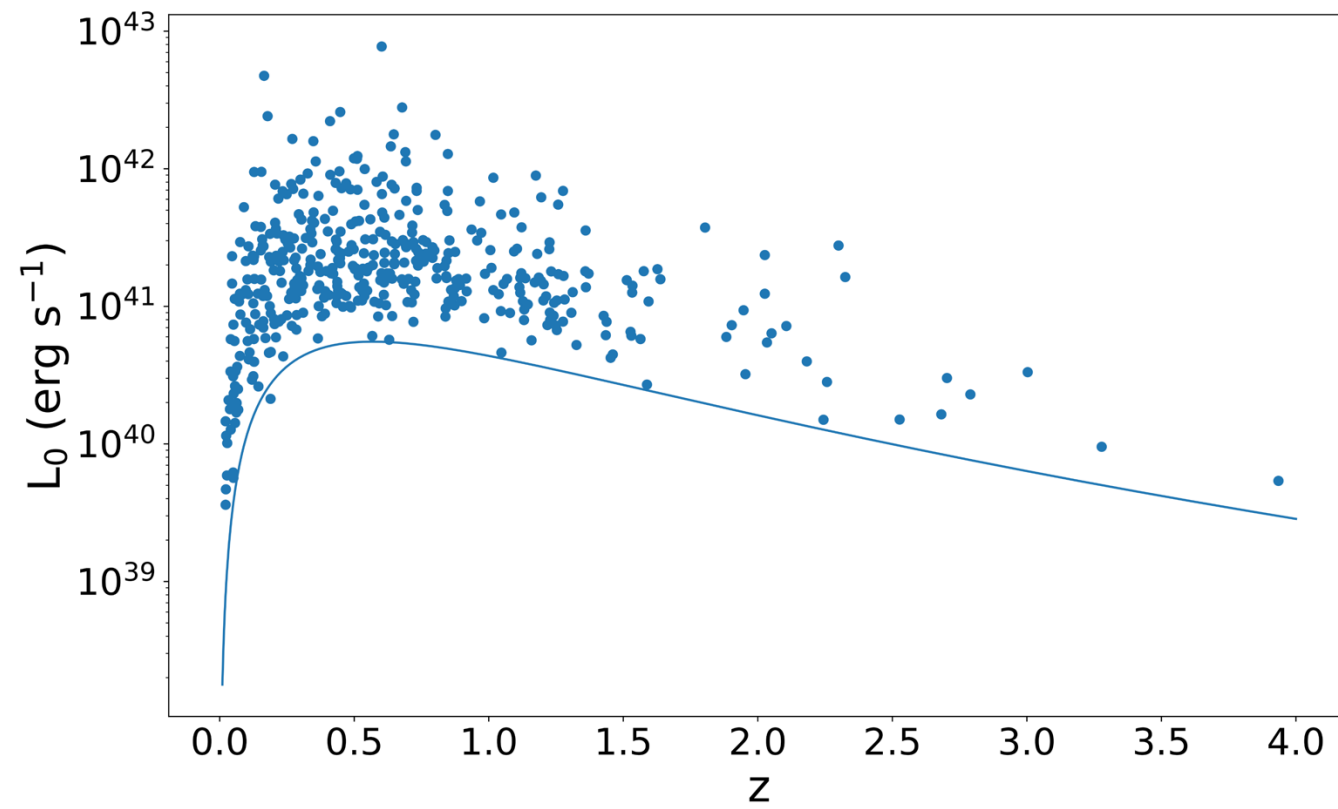


Luminosity-redshift distribution

The degeneracy between the luminosity function and formation rate.

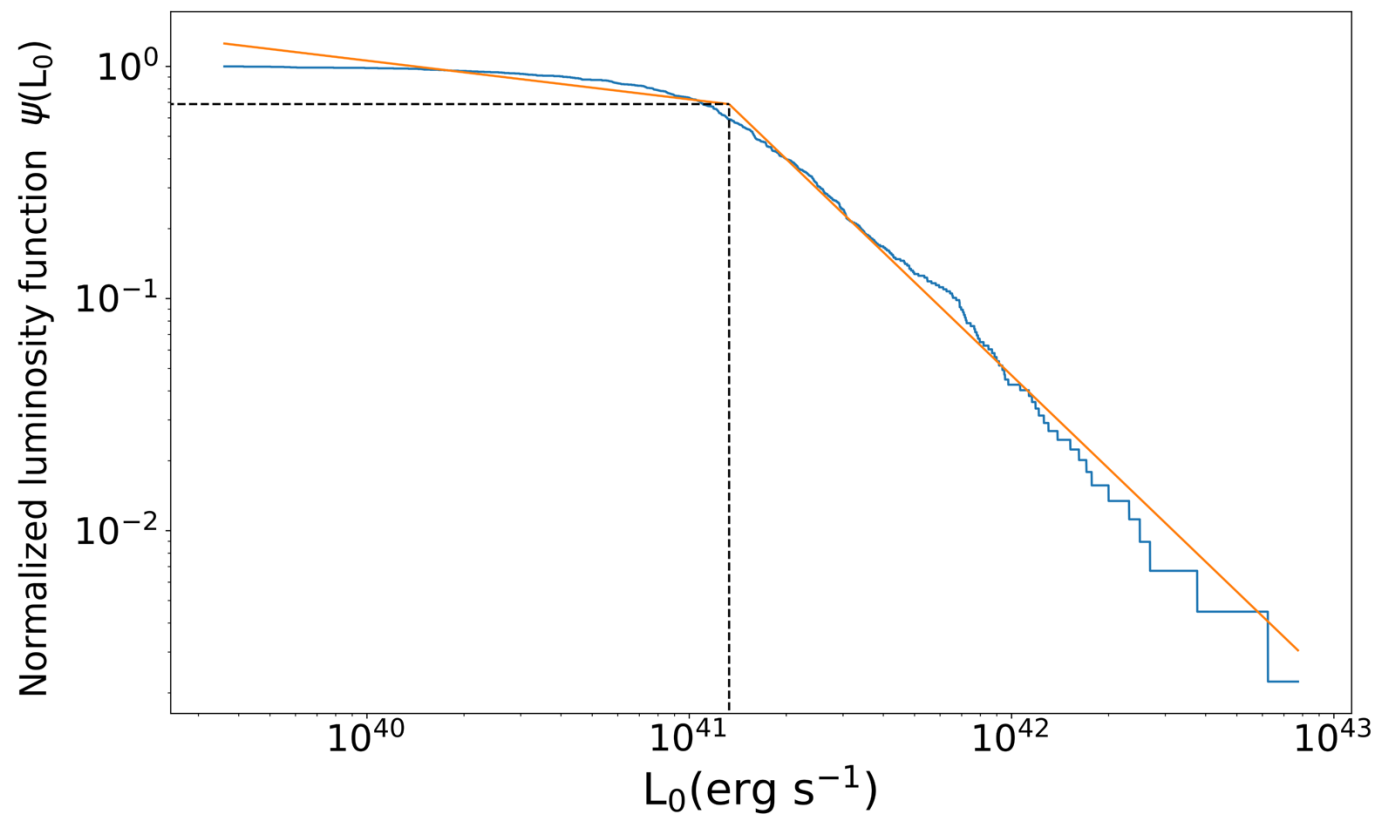
$$\begin{cases} \Psi(L, z) = \psi_z(L)\phi(z) = \psi(L_0)\phi(z) \\ L_0 = L / g(z), \quad g(z) = (1+z)^k \end{cases}$$

Result:



Non-evolving luminosity-redshift distribution

Result:

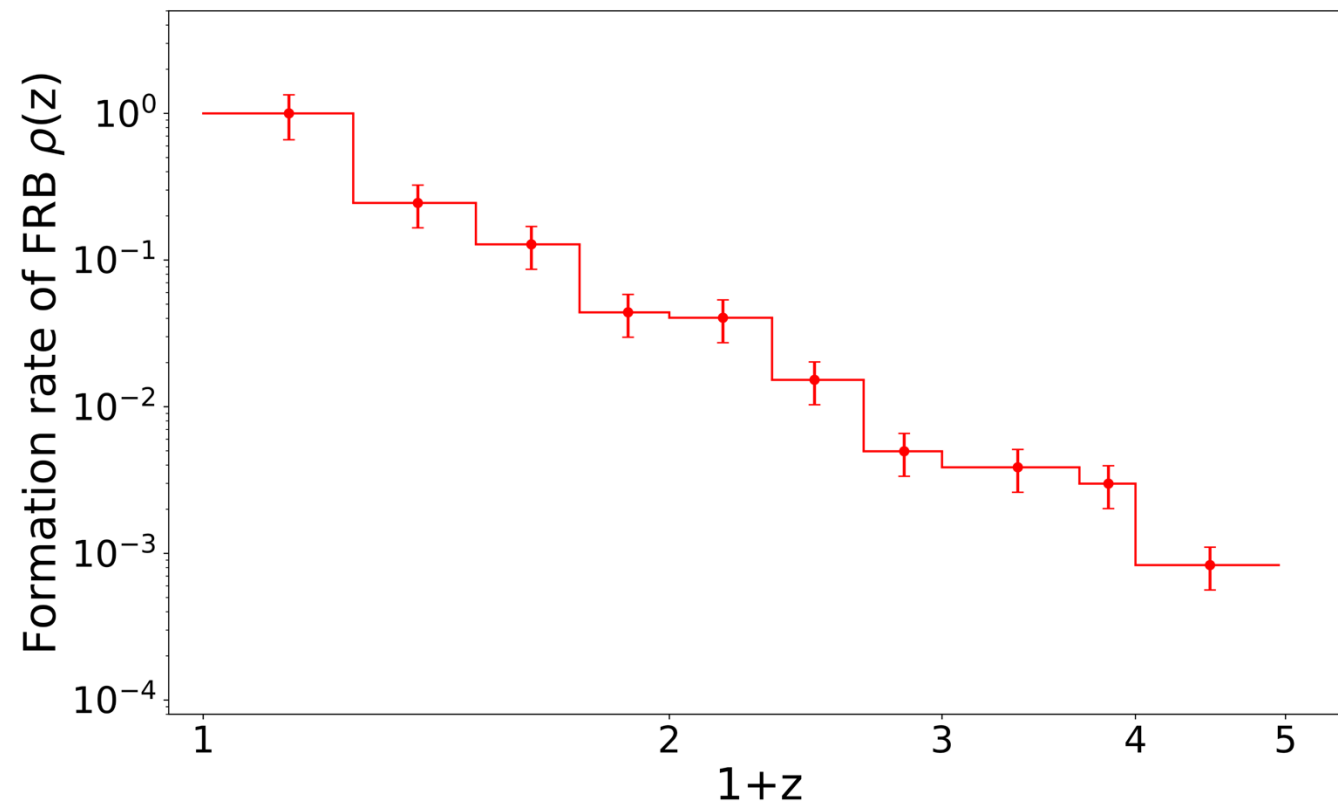


$$\psi(L_0) \propto \begin{cases} L_0^{-0.17 \pm 0.01} & L_0 < L_0^b \\ L_0^{-1.33 \pm 0.01} & L_0 > L_0^b \end{cases}$$

$$L_0^b = 1.33 \times 10^{41} \text{ erg s}^{-1}$$

Cumulative luminosity function

Result:

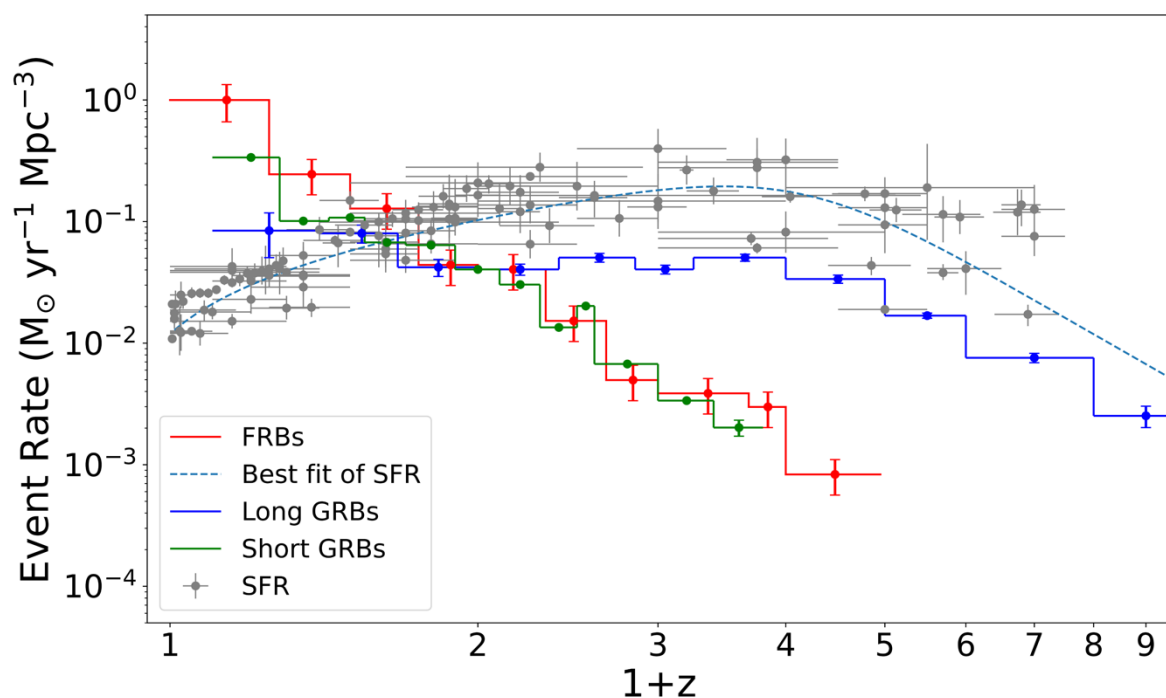


$$\rho(z) \propto (1+z)^{-4.9 \pm 0.3}$$

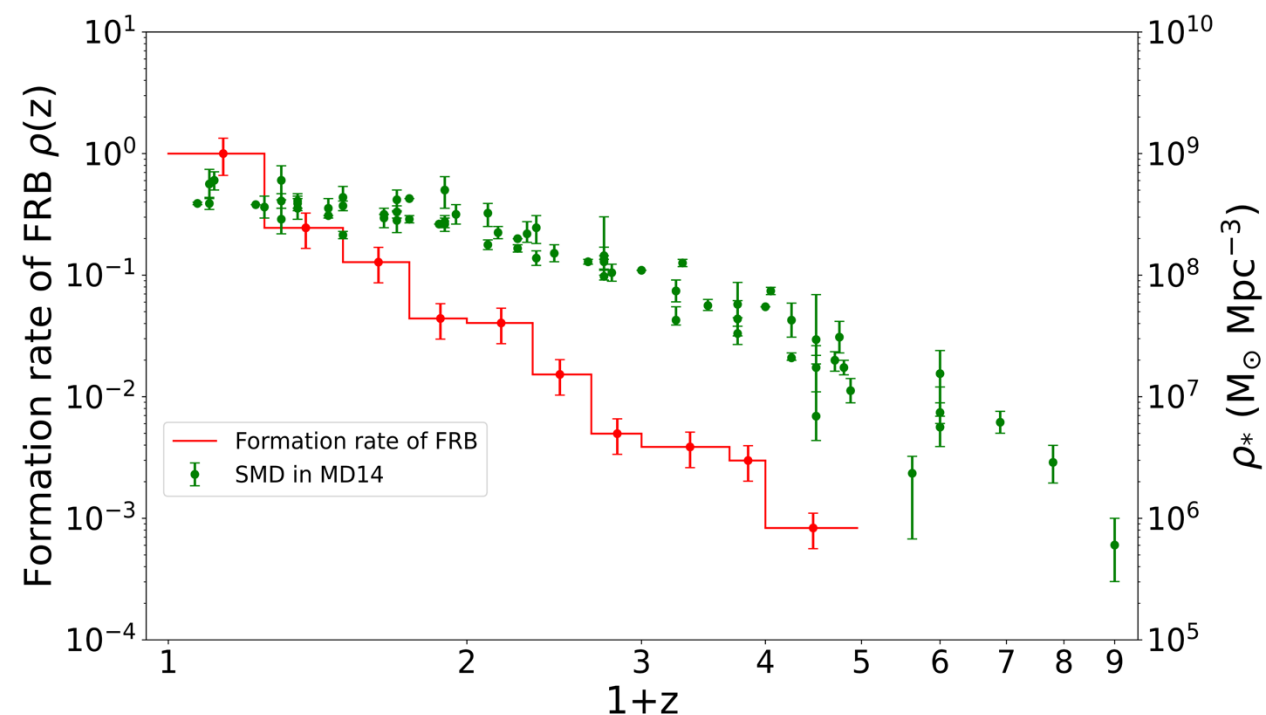
Comoving formation rate $\rho(z)$ of FRBs.

Result:

1. FRB rate deviates from SFR, with a similar redshift dependence as short GRBs.
2. They show similar decreasing trend at $z > 1.0$.



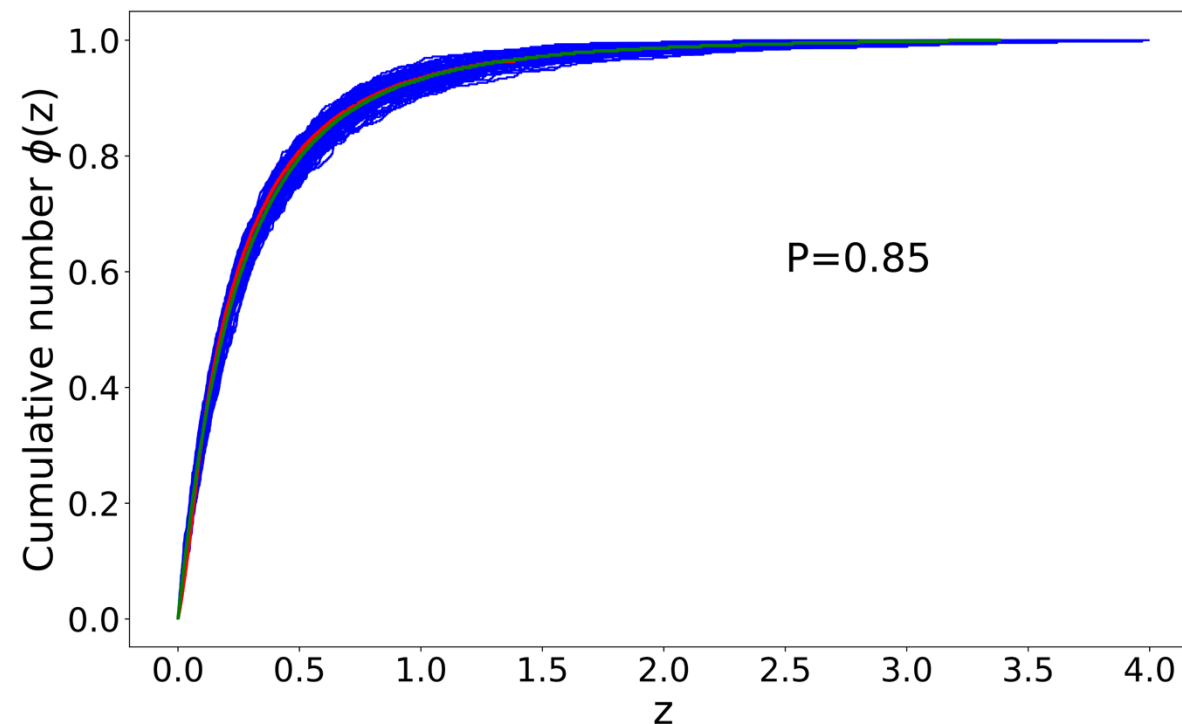
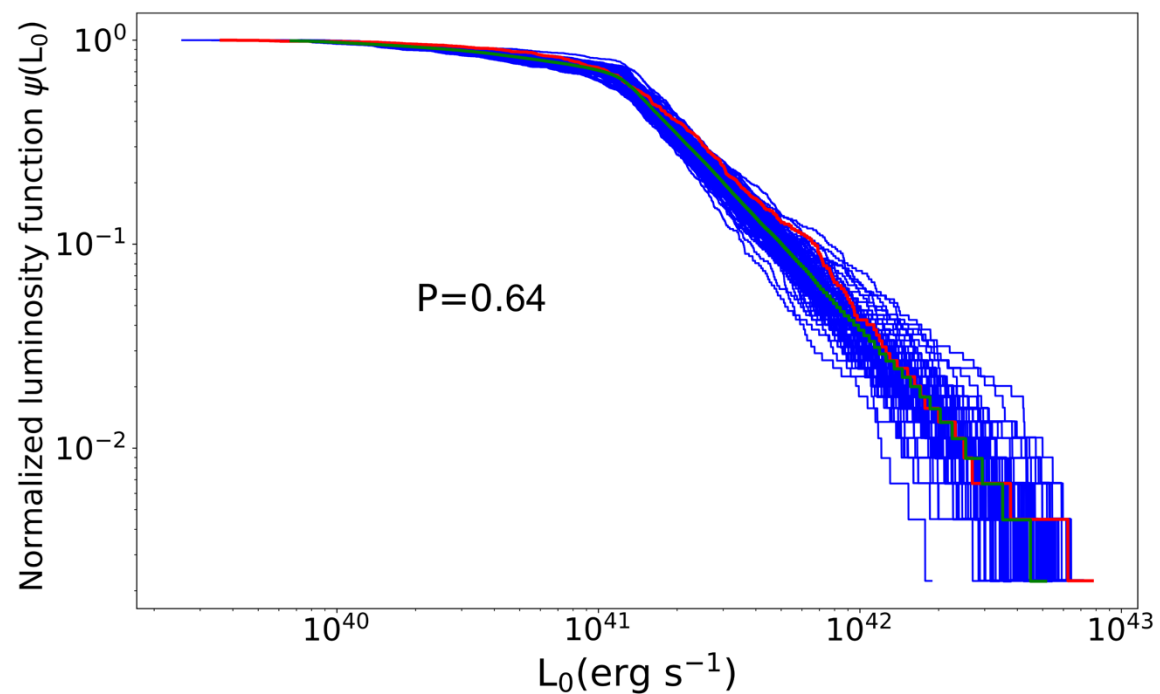
FRB formation rate vs other events.



FRB formation rate vs the observed SMD.

Result:

Monte Carlo simulation



Comparison between the formation rate of FRBs and other events.

Thanks for your listening!

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